Electrodynamics and Yang-Mills theory in $SO(2,3)_{\star}$ model of Noncommutative Gravity

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GRAVITY and STRING THEORY: NEW IDEAS FOR UNSOLVED PROBLEMS III

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Short overview

- **1** Canonical deformation of continuous structure of spacetime; algebra of coordinates: $[\hat{x}^{\mu}, \hat{x}^{\nu}] = i\theta^{\mu\nu}$; representation of the algebra via NC Moyal-Weyl \star -product.
- **2** SO(2,3) gauge theory of gravity: unification of spin-connection and vierbein; gravity emerges only after gauge SB $SO(2,3) \rightarrow SO(1,3)$
- NC deformation of GR:

$$\widehat{S} = -\frac{1}{16\pi G_N} \int d^4x \sqrt{-g} \left(R + \frac{\theta^{\alpha\beta}\theta^{\gamma\delta}}{2l^4} R_{\alpha\beta\gamma\delta} - \frac{15}{16l^4} T_{\alpha\beta}^{\rho} T_{\gamma\delta\rho} + \ldots \right) \right)$$

- **9** NC deformation of Minkowski space; interpretation of θ -constant noncommutativity; *Fermi inertial coordinates*
- ullet Dirac spinors are incorporated; NC corrections are linear in $\theta^{\alpha\beta}$ and they modify the propagator and dispersion relation of a free electron; helicity dependent energy levels; birefringence effect

- D. Gočanin and V. Radovanović, Dirac field and gravity in NC SO(2,3)* model, Eur. Phys. J. C, 78 (2018) 195.
- M. Dimitrijević-Ćirić, D. Gočanin, N. Konjik and V. Radovanović, Noncommutative Electrodynamics from SO(2,3)_{*} Model of Noncommutative Gravity, Eur. Phys. J. C, 78 (2018) 548.
- M. Dimitrijević-Ćirić, D. Gočanin, N. Konjik and V. Radovanović, Yang-Mills Theory in the SO(2,3)_⋆ Model of Noncommutative Gravity (under revision)

Next step: gauge fields in SO(2,3) framework

Unifying gravitational field with $Maxwell\ U(1)$ gauge field and $Yang-Mills\ SU(N)$ gauge field within SO(2,3) model.

Gauge group $SO(2,3)\times SU(N)$. Generators T_I of SU(N) group are hermitian, traceless and they satisfy the (anti)commutation relations: $[T_I,T_J]=if_{IJK}T_K$ and $\{T_I,T_J\}=d_{IJK}T_K$; normalization $Tr(T_IT_J)=\delta_{IJ}$.

Unifying gauge potential:

$$\Omega_{\mu} = \frac{1}{2} \omega_{\mu}^{AB} M_{AB} \otimes \mathbb{I} + \mathbb{I} \otimes A_{\mu}^{I} T_{I} \equiv \frac{1}{4} \omega_{\mu}^{ab} \sigma_{ab} - \frac{1}{2I} \mathbf{e}_{\mu}^{a} \gamma_{a} + A_{\mu}^{I} T_{I}$$

Field strength:

$$\mathbb{F}_{\mu\nu} = \partial_{\mu}\Omega_{\nu} - \partial_{\nu}\Omega_{\mu} - i[\Omega_{\mu}, \Omega_{\nu}] = \frac{1}{2}F_{\mu\nu}^{AB}M_{AB} \otimes \mathbb{I} + \mathbb{I} \otimes \mathcal{F}^{I}T_{I}$$

$$F_{\mu\nu}^{ab} = R_{\mu\nu}^{ab} - rac{2}{l^2} e_{(\mu}^a e_{
u)}^b, \hspace{0.5cm} F_{\mu
u}^{a5} = rac{1}{l} rac{T_a}{\mu
u} \hspace{0.5cm} \mathcal{F}_{\mu
u}^l = 2 \partial_{(\mu} A_{
u)}^l + g f^{IJK} A^J A^K$$



Commutative model

Commutative (un-deformed) action:

$$S_A = -rac{1}{16l} Tr \int d^4x \; arepsilon^{\mu
u
ho\sigma} \Big\{ f \mathbb{F}_{\mu
u} D_
ho\phi D_\sigma\phi\phi + rac{i}{3!} f^2 D_\mu\phi D_
u\phi D_
ho\phi D_\sigma\phi\phi \Big\} + h.c.$$

- lacktriangledown action is hermitian and possesses $SO(2,3) \times SU(N)$ gauge symmetry
- ② we don't have metric a priori and thus no Hodge dual operation \rightarrow auxiliary field $f=\frac{1}{2}f^{AB,I}M_{AB}\otimes T_{I}$
- **3** auxiliary field $\phi = \phi^A \Gamma_A$ is used for symmetry breaking: $\phi^a = 0$ and $\phi^5 = I$; covariant derivative after SB: $(D_\mu \phi)^a = e^a_\mu$ and $(D_\mu \phi)^5 = 0$
- it reduces to the canonical kinetic action for SU(N) gauge field in curved spacetime after the SB:

$$S_A = -rac{1}{4}\int d^4x \, \sqrt{-g} \, g^{\mu\rho}g^{
u\sigma}\mathcal{F}^I_{\mu
u}\mathcal{F}^I_{
ho\sigma}$$

NC-deformed action

ordinary commutative product ightarrow noncommutative Moyal-Weyl \star -product

$$\begin{split} \widehat{S}_{A} &= -\frac{1}{16I} T_{r} \int d^{4}x \; \varepsilon^{\mu\nu\rho\sigma} \left\{ \widehat{f} \star \widehat{\mathbb{F}}_{\mu\nu} \star D_{\rho} \widehat{\phi} \star D_{\sigma} \widehat{\phi} \star \widehat{\phi} \right. \\ &+ \frac{i}{6} \widehat{f} \star \widehat{f} \star D_{\mu} \widehat{\phi} \star D_{\nu} \widehat{\phi} \star D_{\rho} \widehat{\phi} \star D_{\sigma} \widehat{\phi} \star \widehat{\phi} \right\} + h.c. \end{split}$$

transformations. Symmetry is noncommutatively preserved!

This action can now be perturbatively expanded in powers of $\theta^{\alpha\beta}$ and the

Deformed action is hermitian and invariant under $SO(2,3)_{\star} \times SU(N)_{\star}$ gauge

This action can now be perturbatively expanded in powers of $\theta^{\alpha\beta}$ and the Seiberg-Witten map insures that, order-by-order, this expansion possesses ordinary $SO(2,3)\times SU(N)$ gauge symmetry of the original undeformed action S_A .

After the symmetry breaking, the leading term in NC perturbative expansion turns out to be linear in $\theta^{\alpha\beta}$:

$$\widehat{S}_A^{(1)} = -rac{ heta^{lphaeta}}{16}\int d^4x\; e\; d_{IJK}\; g^{\mu
ho}g^{
u\sigma}\left(\mathcal{F}_{lphaeta}^I\mathcal{F}_{\mu
u}^J\mathcal{F}_{
ho\sigma}^K - 4\mathcal{F}_{lpha\mu}^I\mathcal{F}_{eta
u}^J\mathcal{F}_{
ho\sigma}^K
ight)$$

Fermions

NC-deformed $SO(2,3)_{\star} \times SU(N)_{\star}$ invariant spinor action before SB:

$$\begin{split} \widehat{S}_{\Psi} &= \frac{i}{12} \int d^4x \; \varepsilon^{\mu\nu\rho\sigma} \; \left[\widehat{\bar{\Psi}} \star \left(D_{\mu} \widehat{\phi} \right) \star \left(D_{\nu} \widehat{\phi} \right) \star \left(D_{\rho} \widehat{\phi} \right) \star \left(D_{\sigma} \widehat{\Psi} \right) \right. \\ & \left. - \left(D_{\sigma} \widehat{\bar{\Psi}} \right) \star \left(D_{\mu} \widehat{\phi} \right) \star \left(D_{\nu} \widehat{\phi} \right) \star \left(D_{\rho} \widehat{\phi} \right) \star \widehat{\Psi} \right] + \widehat{S}_{m} \end{split}$$

Linear NC correction of Yang-Mills theory in curved spacetime after the SB:

$$\widehat{S}_{\Psi}^{(1)} = \theta^{\alpha\beta} \int d^4x \ e \ \overline{\Psi} \Big(\mathcal{A}_{\alpha\beta}^{\rho\sigma} \mathcal{D}_{\rho} \mathcal{D}_{\sigma} + \mathcal{B}_{\alpha\beta}^{\sigma} \mathcal{D}_{\sigma} + \mathcal{C}_{\alpha} \mathcal{D}_{\beta} + \mathcal{D}_{\alpha\beta} \Big) \Psi$$

New terms: $\bar{\Psi}\sigma_{\alpha}{}^{\sigma}\mathcal{D}_{\beta}\mathcal{D}_{\sigma}\Psi$, $\bar{\Psi}R_{\alpha\beta}{}^{\rho\sigma}\gamma_{\rho}\mathcal{D}_{\sigma}\Psi$, $\bar{\Psi}T_{\alpha\beta}{}^{\sigma}\mathcal{D}_{\sigma}\Psi$, $\bar{\Psi}\sigma_{\alpha\beta}\Psi$, ...

 $SO(1,3) \times SU(N)$ covariant derivative:

$$(\mathcal{D}_{\sigma}\Psi)_{i} = \nabla_{\sigma}\psi_{i} - i\mathcal{A}_{\sigma}^{I}(T_{I})_{ij}\psi_{j}$$



NC Yang-Mills theory in Minkowski space

Residual effects of noncommutativity in flat spacetime produce new types of coupling in Yang-Mills theory:

$$\begin{split} \widehat{S}_{\textit{flat}} &= \int d^4x \left\{ i \bar{\Psi} \mathcal{D} \Psi - \frac{1}{4} g^{\mu\rho} g^{\nu\sigma} \mathcal{F}^I_{\mu\nu} \mathcal{F}^I_{\rho\sigma} \right. \\ &+ \theta^{\alpha\beta} \left[-\frac{1}{2I} \bar{\Psi} \sigma_{\alpha}{}^{\sigma} \mathcal{D}_{\beta} \mathcal{D}_{\sigma} \Psi + \frac{7i}{24I^2} \varepsilon_{\alpha\beta}{}^{\rho\sigma} \bar{\Psi} \gamma_{\rho} \gamma_5 \mathcal{D}_{\sigma} \Psi + \frac{3i}{4} \bar{\Psi} \mathcal{F}_{\alpha\beta} \mathcal{D} \Psi \right. \\ &- \frac{i}{2} \bar{\Psi} \mathcal{F}_{\alpha\sigma} \gamma^{\sigma} \mathcal{D}_{\beta} \Psi - \frac{1}{6I^3} \bar{\Psi} \sigma_{\alpha\beta} \Psi + \frac{1}{4I} \bar{\Psi} \mathcal{F}_{\alpha\beta} \Psi \\ &- \frac{1}{16} d_{IJK} g^{\mu\rho} g^{\nu\sigma} \left(\mathcal{F}^I_{\alpha\beta} \mathcal{F}^J_{\mu\nu} \mathcal{F}^K_{\rho\sigma} - 4 \mathcal{F}^I_{\alpha\mu} \mathcal{F}^J_{\beta\nu} \mathcal{F}^K_{\rho\sigma} \right) \right] \right\} \end{split}$$

Dispersion relation for "free" electrons

NC Dirac equation: $(i\partial \!\!\!/ - m + \not \!\!\!/ + \theta^{\alpha\beta} \mathcal{M}_{\alpha\beta}) \psi = 0$

For "free" electron, $[H,\mathbf{p}]=0\Rightarrow$ plane wave ansatz $\psi(x)=u(\mathbf{p})e^{-ip\cdot x}$

For $\mathbf{p}=(0,0,p_z)$ and $[\hat{\chi}^1,\hat{\chi}^2]=i\theta^{12},\ \theta^{12}\equiv\theta\neq0$, we have four independent energy functions:

$$E_{1,2} = E_{\mathbf{p}} \mp \left[\frac{m^2}{12I^2} - \frac{m}{3I^3} \right] \frac{\theta}{E_{\mathbf{p}}} + \mathcal{O}(\theta^2)$$

$$E_{3,4} = -E_{\mathbf{p}} \pm \left[\frac{m^2}{12I^2} - \frac{m}{3I^3} \right] \frac{\theta}{E_{\mathbf{p}}} + \mathcal{O}(\theta^2)$$

with classical energy $E_{\mathbf{p}} = \sqrt{m^2 + p_z^2}$. \Rightarrow **Zeeman effect!**

Important observation: Electron's energy depends on its *helicity*! Background NC space behaves as a *birefringent* medium for electrons propagating in it.

Electron in magnetic field - NC Landau levels

Special case: electron in magnetic field $\mathbf{B} = B\mathbf{e}_z$

NC-deformed energy levels:

$$E_{n,s} = \sqrt{p_z^2 + m^2 + (2n + s + 1)B}$$

$$- \theta \left[\frac{mM^3s}{E_{n,s}^{(0)}} + \frac{M^3Bs}{E_{n,s}^{(0)}(E_{n,s}^{(0)} + m)} (2n + s + 1) - \frac{B^2}{2E_{n,s}^{(0)}} (2n + s + 1) \right]$$

Non-relativistic limit, $B \ll m^2$, for an electron confined in NC plane:

$$E_{n,s} = m_{eff} + \frac{2n+s+1}{2m} B_{eff} - \frac{(2n+s+1)^2}{8m^3} B_{eff}^2$$

$$m_{eff} = m - s\theta M^3, \ B_{eff} = (B + \theta B^2), \ M^3 = \left[\frac{m}{12l^2} - \frac{1}{3l^3}\right].$$

Inverse of the canonical deformation parameter θ plays the role of an effective magnetic field $b=\theta^{-1}$.

NC correction to magnetic dipole moment of an electron:

$$\mu_{n,s} = -\frac{\partial E_{n,s}}{\partial B} = -\mu_B (2n+1+s)(1+\frac{\theta B}{B})$$

Further development of $SO(2,3)_{\star}$ model

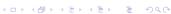
Renormalizability: Minimal NC Electrodynamics,

$$\widehat{S} = \int d^4x \ \widehat{\overline{\psi}} \star (i\gamma^\mu D_\mu - m)\widehat{\psi} - \frac{1}{4} \int d^4x \ \widehat{F}_{\mu\nu} \star \widehat{F}^{\mu\nu}$$

is not renormalizabile. Minimal NC Yang-Mills is one-loop renormalizabile to first order in θ .

- NC Standard Model: introducing (complex) scalars
- **NC SUGRA**: Introduce local SUSY; spin-3/2 gravitino field; orthosymplectic OSp(1,4) group;

- NC quantum Hall effect: How noncommutativity influences the phenomenology of the Hall effect?
- Quark-gluon plasma in NC spacetime: How noncommutativity influences the phenomenology of quark-gluon plasma in the early stages of the Universe?



The End